

Modelling Radiative Line Transfer in Maser Environments

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Overview

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The Rate Equation

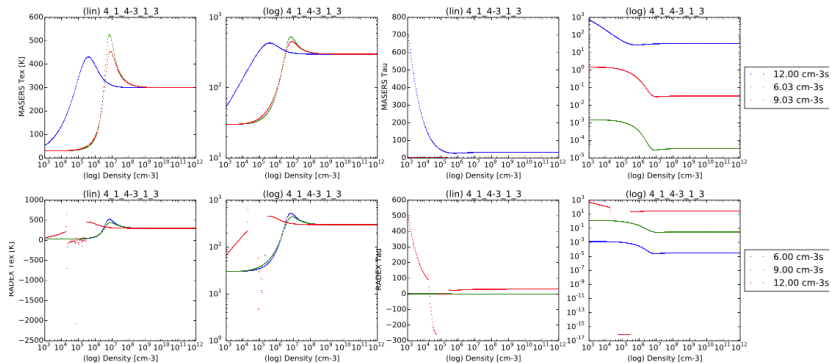
$$\begin{aligned} \frac{dn_i}{dt} = & - \sum_{j < i} \left\{ A_{ij} \beta_{ij} [n_i + X_{ij}(n_i - n_j)] + C_{ij} \left[n_i - n_j \exp\left(\frac{-h\nu_{ij}}{kT}\right) \right] \right\} \\ & + \sum_{j > i} \frac{g_j}{g_i} \left\{ A_{ji} \beta_{ji} [n_j + X_{ji}(n_j - n_i)] + C_{ji} \left[n_j - n_i \exp\left(\frac{-h\nu_{ji}}{kT}\right) \right] \right\} \end{aligned}$$

$$R_{ij} = \begin{cases} A_{ji} \beta_{ji} (1 + X_{ji}) & i < j \\ A_{ij} \beta_{ij} \left(\frac{g_i}{g_j}\right) X_{ij} & j > i \\ - \sum_{i \neq j} R_{ji} & j = i \end{cases}$$

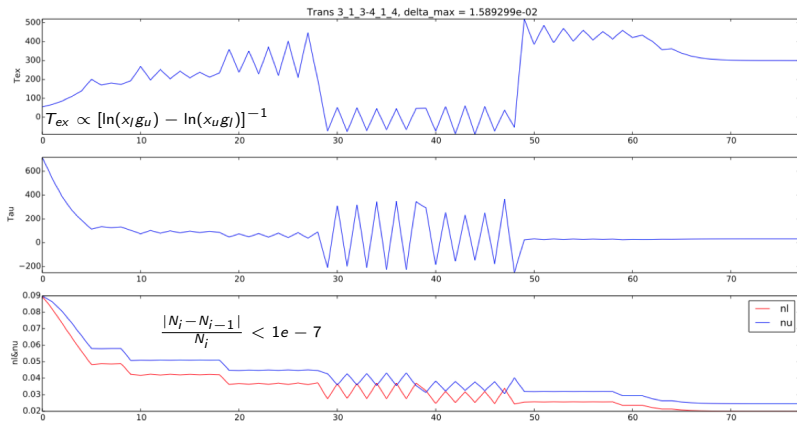
$$\tilde{C}_{ij} = \begin{cases} C_{ji} & j > i \\ - \sum_{i \neq j} C_{ij} & j = i \end{cases}$$

$$(\mathbf{R} + \tilde{\mathbf{C}})\mathbf{n} = \mathbf{b}$$

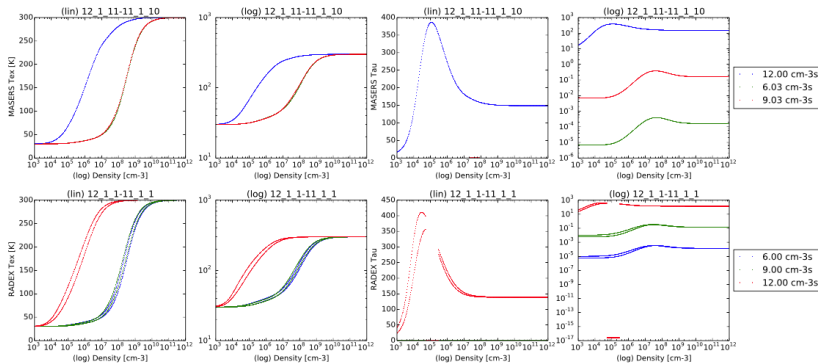
Numerical Issues



Convergence Criteria

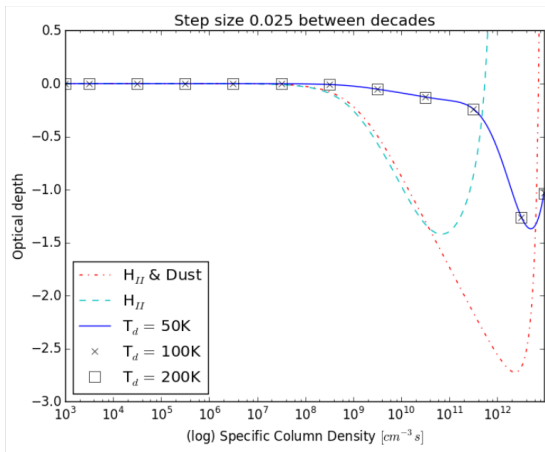


Relaxation

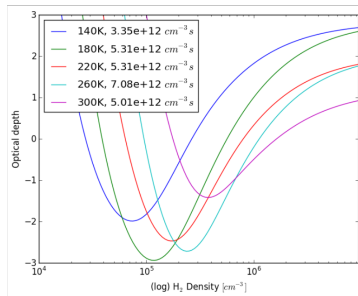
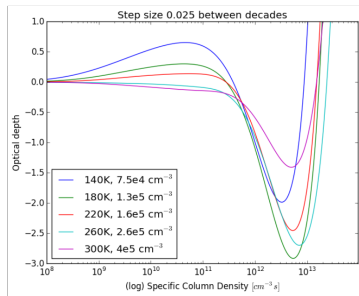


G37.55 + 0.20 4.8 GHz H_2CO Maser

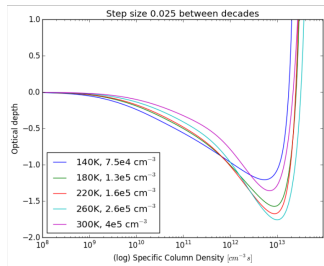
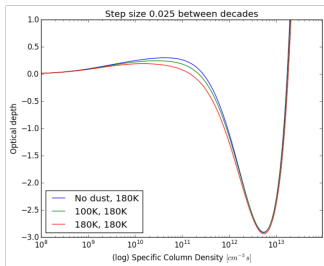
$$X = \beta_{ij} W_{HII} N_{HII} + \beta_{ij} W_d N_d$$



Collisional Excitation



Radiative Excitation



Collision & dust radiation: $w_d = 1$, $n_{H_2} = 1.3e5 \text{ cm}^{-3}$

Collision & free-free emission: $w_{HII} = 1$, $EM = 1e9 \text{ pc cm}^{-6}$

Line Overlap

$$R_{ij}^a = \begin{cases} A_{ji}\beta_{ji}^a(1 + X_{ji}) & i < j \\ A_{ij}\beta_{ij}^a \left(\frac{g_i}{g_j}\right) X_{ij} & j > i \\ -\sum_{i \neq j} R_{ji} & j = i \end{cases}, \quad R_{ij}^b = \begin{cases} A_{ji}\beta_{ji}^b & i < j \\ A_{ij}\beta_{ij}^b & j > i \\ 0 & j = i \end{cases},$$

$$R_{ij}^{ab} = \begin{cases} A_{ji}\beta_{ji}^{ab}(x_a + X_{ji}) & i < j \\ A_{ij}\beta_{ij}^{ab} \left(\frac{g_i}{g_j}\right) X_{ij} & j > i \\ -\sum_{i \neq j} R_{ji} & j = i \end{cases},$$

$R = R^a + R^b = R^{ab}$ where $\beta^a = [1 - (1 - f_a)(1 - \beta_{\tau_a})]$,
 $\beta^{ab} = f_a(1 - \beta_{\tau_a + \tau_b})$, $\beta^b = f_b x_a(1 - \beta_{\tau_a + \tau_b})$ and $x_a = \frac{\kappa_a}{\kappa_{ab}}$ and $f_a = \frac{\Delta_{ab}}{\Delta_a}$
the absorption and overlap ratios respectively.

References



Elitzur M 1992 *Astronomical Masers* (Netherlands: Kluwer Academic Publishers)



Schöier F L, van der Tak F F S, van Dishoeck E F and Black J H An atomic and molecular database for analysis of submillimetre line observations 2005 *A&A* **432** 369-79



Van der Tak F F S, Black J H, Schöier F L, Jansen D J and van Dishoeck E F A computer program for fast non-LTE analysis of interstellar line spectra. With diagnostic plots to interpret observed line intensity ratios 2007 *A&A* **468** 627-35



Wiesenfeld L and Faure A Rotational quenching of H_2CO by molecular hydrogen: cross-sections, rates and pressure broadening 2013



Van der Walt D J Pumping of the 4.8 GHz H_2CO masers and its implications for the periodic masers in G37.55+0.20 2014 *A&A* **562** A68



Elitzur M and Netzer H Line fluorescence in astrophysics 1885 *ApJ* **291** 464-67