



Fast Scheme for Approximating an Offset PSF Response

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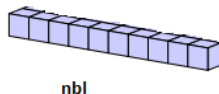
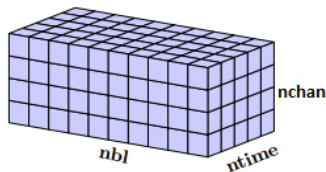
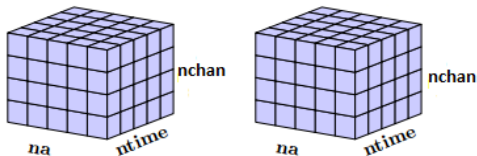
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SAIP

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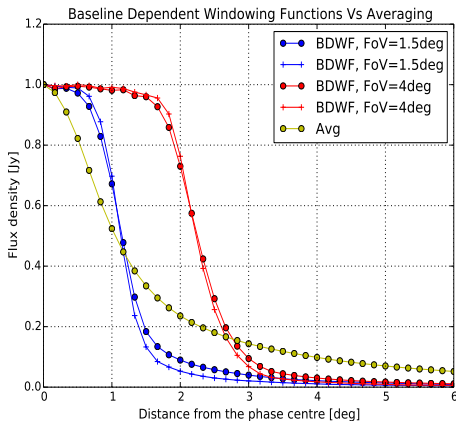
Problem statement: decorrelation

- Visibility averaging
- Time and bandwidth decorrelation
- Smearing = loss of amplitude + distortion



Problem statement: decorrelation

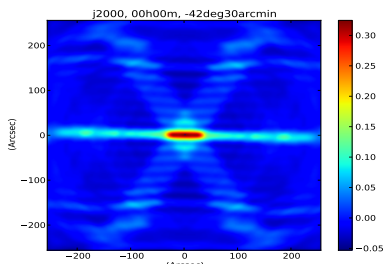
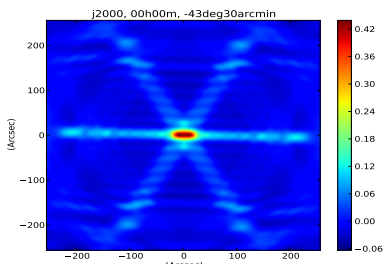
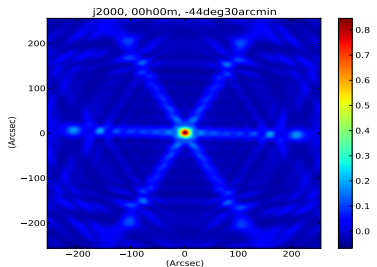
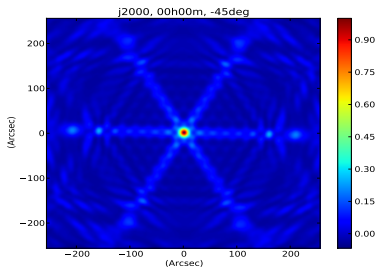
- Visibility averaging
- Time and bandwidth smearing
- Smearing = loss of amplitude + distortion



JVLA C, 100s integration and 10MHz channels width

Averaging effects or smearing

- Phase center RA=0, Dec=-45 deg



Problem solving: PSF slow derivation

- interferometric visibility

$$V(\mathbf{b}) = \int_{\Omega} I_{\nu}(\mathbf{s}) e^{-2\pi i \mathbf{b} \mathbf{s}} d\Omega, \quad (1)$$

- $t\nu$ -space: re-sampling time Δt and frequency $\Delta \nu$, t_0 and ν_0 their centre respectively

$$V(t_0, \nu_0) = \iint_{\Delta t \Delta \nu} W_{t_0, \nu_0}(t, \nu) \left[\int_{\Omega} I_{\nu}(\mathbf{s}) e^{-2\pi i \mathbf{b} \mathbf{s}} d\Omega \right] dt d\nu \quad (2)$$

$$= \iint_{u_{t\nu} v_{t\nu}} W_{\mathbf{b}_0}(\mathbf{b}) \left[\int_{\Omega} I_{\nu}(\mathbf{s}) e^{-2\pi i \mathbf{b} \mathbf{s}} d\Omega \right] du_{t\nu} dv_{t\nu} \quad (3)$$

- Generalized weighting and windowing function

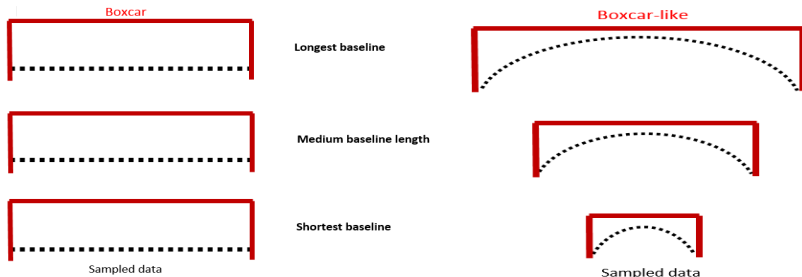
$$W_{t_0, \nu_0}(t, \nu) = \Pi(t - t_0, \nu - \nu_0) W(t - t_0, \nu - \nu_0) \quad (4)$$

tv -space and uv -space

- Boxcar and boxcar-like

$$\Pi(t - t_0, \nu - \nu_0) = \begin{cases} \frac{1}{\Delta t \Delta \nu}, & |t| \leq \Delta t/2, \quad |\nu| \leq \Delta \nu/2 \\ 0, & \text{otherwise,} \end{cases} \quad (5)$$

$$\Pi(\mathbf{b} - \mathbf{b}_0) = \begin{cases} \frac{1}{\Delta u_{t\nu} \Delta v_{t\nu}}, & |u_{t\nu}| \leq \Delta u_{t\nu}/2, \quad |v_{t\nu}| \leq \Delta v_{t\nu}/2 \\ 0, & \text{otherwise,} \end{cases} \quad (6)$$



$$V_{\mathbf{b}_0}(\mathbf{x}) = \iint_{u_{t\nu} v_{t\nu}} W_{\mathbf{b}_0}(\mathbf{b}) e^{-2\pi i(u_{t\nu}l + v_{t\nu}m + w_{t\nu}(n-1))} du_{t\nu} dv_{t\nu}. \quad (7)$$

- source coordinates, l , m and n . Average baseline vector, \mathbf{x}

$$V_{\mathbf{b}_0}(\mathbf{x}) = \left(\tilde{G}_{\mathbf{b}} \circ [C_{\mathbf{b}_0} \tilde{\Pi}_{\mathbf{b}}] \circ [C_{\mathbf{b}_0} \tilde{W}_{\mathbf{b}}] \right)_{(l,m)} \quad (8)$$

$$C_{\mathbf{b}_0}(l, m) = e^{-2\pi i(u_{t_0\nu_0}l + v_{t_0\nu_0}m)}, \quad G_{\mathbf{b}}(l, m) = e^{-2\pi i[w_{t\nu}(n-1)]} \quad (9)$$

$$V_{b_0}(\mathbf{x}) = \left(\tilde{G}_b \circ [C_{b_0} \tilde{\Pi}_b] \circ [C_{b_0} \tilde{W}_b] \right)_{(l,m)} \quad (10)$$

Brute force: (CPU, Parallel computing, GPU,...)

- Slow derivation, FFT, sky map of N^2 cells, $[PSF]_{l,m}$ requires $\sim MN \log_2 N$ multiplications; $\sim N^2 \log_2 N$.
- Slow derivation, FFT, sky map of N_s sources, $\sim N^4 \log_2 N$ multiplications for all $[PSF]_{l,m}$.

Linear Algebra and Approximation: computationally cheaper, but an approximation

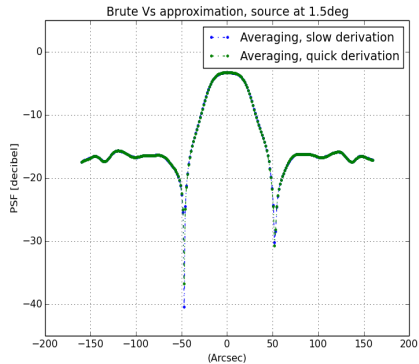
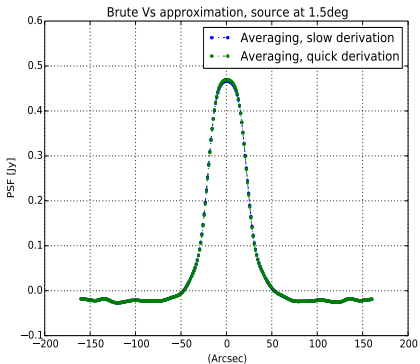
- Quick derivation, FFT, sky map of N_s sources, $\sim N^2 \log_2 N$ multiplications for all $[PSF]_{l,m}$.

- Approximate the offset PSF (Pseudo PSF) visibilities by the centre bin

$$V_{\mathbf{b}_0}(\mathbf{x}) \simeq \left(\tilde{G}_{\mathbf{b}_0} \circ \left[C_{\mathbf{b}_0} \tilde{\Pi}_{\mathbf{b}_0} \right] \circ \left[C_{\mathbf{b}_0} \tilde{W}_{\mathbf{b}_0} \right] \right)_{(l,m)} \quad (11)$$

- Requires $\sim N \log_2 N$ multiplications

Results



Results

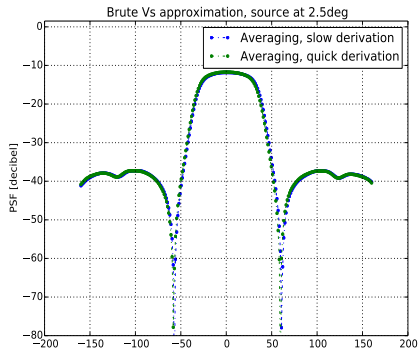
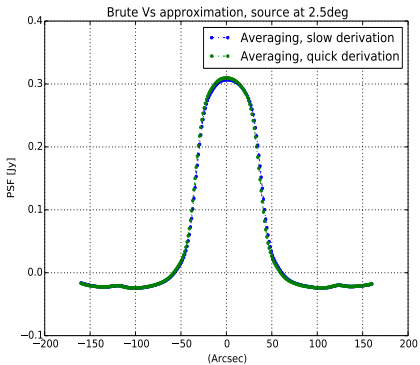


Image plane approximation

- Approximate the offset PSF from the PSF at the Phase centre (Nominal PSF)
- Find a function H and the variation in time/frequency $\Delta\Psi$ and $\Delta\Phi$

$$PSF_{l,m} \simeq \left[H\left(\frac{\Delta\Psi}{2}\right) H\left(\frac{\Delta\Phi}{2}\right) \right] \circ PSF_{l_0,m_0} \quad (12)$$

Algorithm 1 Construct Maxtrices, $\Delta\Psi$ and $\Delta\Phi$

```
1: procedure VARIATION IN TIME AND FREQUENCY
2:    $FoV = N_{pix} \Delta l \Delta m // N_{pix}$  the number of pixels
3:    $\Delta u = \Delta v = 1/FoV$ 
4:    $v_0 = -(N_{pix} - 1) \Delta v / 2$ 
5:    $h = (2v_0) / (N_{pix} - 1)$ ,  $cst = \frac{-\pi}{432 \times 10^2}$ 
6:   for  $i$  from 1 to  $N_{pix}$  do
7:      $u_0 = -(N_{pix} - 1) \Delta u / 2$ 
8:     for  $j$  from 1 to  $N_{pix}$  do
9:        $\theta = \arctan(u_0/v_0)$ 
10:       $du_{i,j} = cst \sqrt{u_0^2 + v_0^2} \sin \theta$ 
11:       $dv_{i,j} = cst \sqrt{u_0^2 + v_0^2} \cos \theta$ 
12:       $\Delta\Psi_{i,j} = \pi(du_{i,j}l + dv_{i,j}m) \Delta t$ 
13:       $\Delta\Phi_{i,j} = \pi \frac{\Delta\nu}{\nu} \sqrt{l^2 + m^2} \sqrt{u_0^2 + v_0^2}$ 
14:       $u_0 = u_0 + h$ 
15:    end for
16:     $v_0 = v_0 + h$ 
17:  end for
18: end procedure
```

$$K = \text{sinc}\left(\frac{\Delta\Psi}{2}\right)\text{sinc}\left(\frac{\Delta\Phi}{2}\right)$$

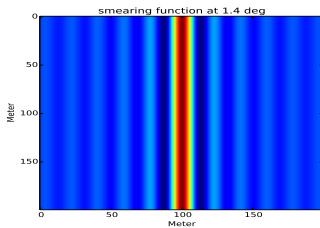
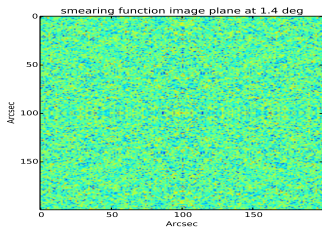
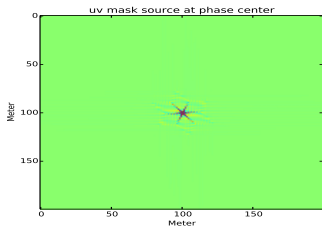
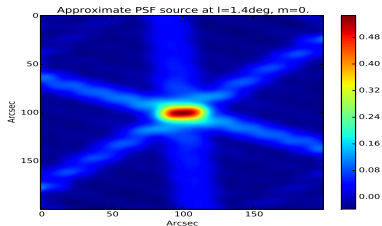
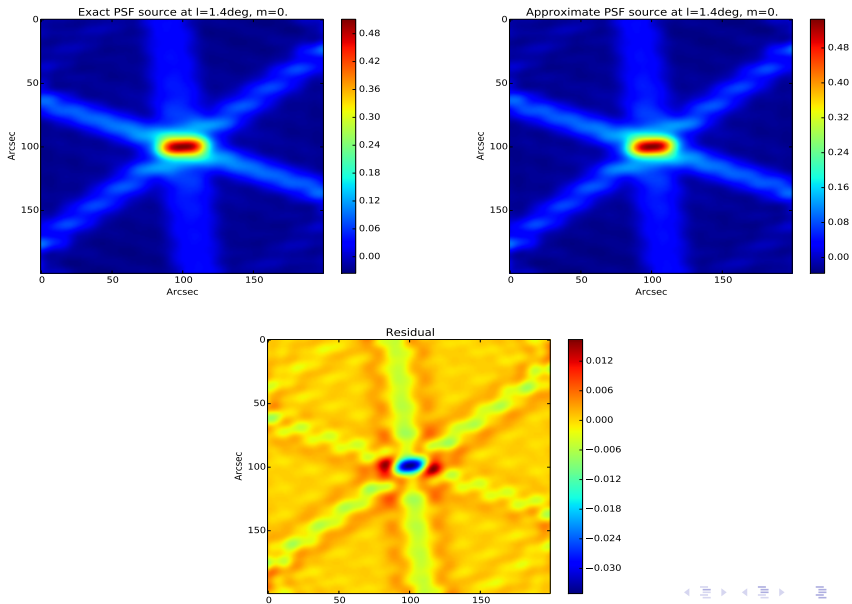
 \tilde{K}

 $\widetilde{PSF}_{l_0, m_0}$

 $\widetilde{PSF}_{l, m}$


Image plane approximation: results



uv-plane approximation

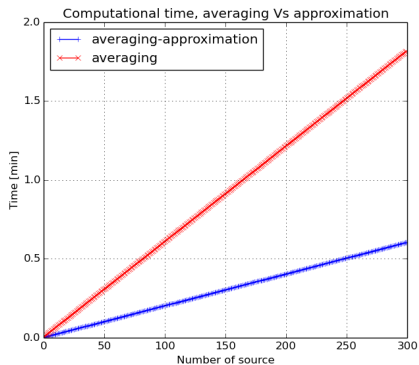
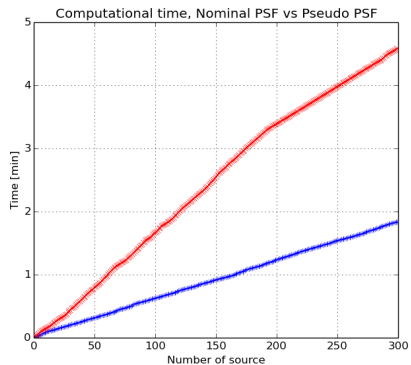
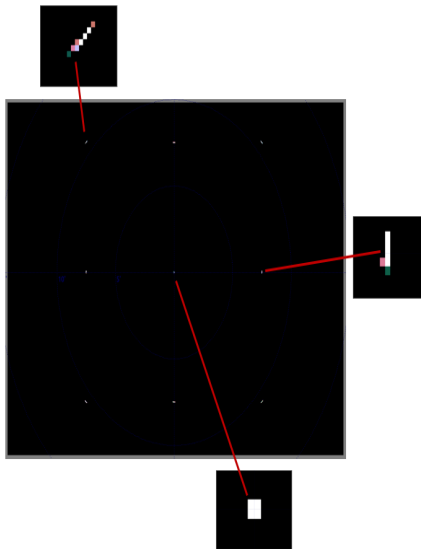


Image plane Approximation

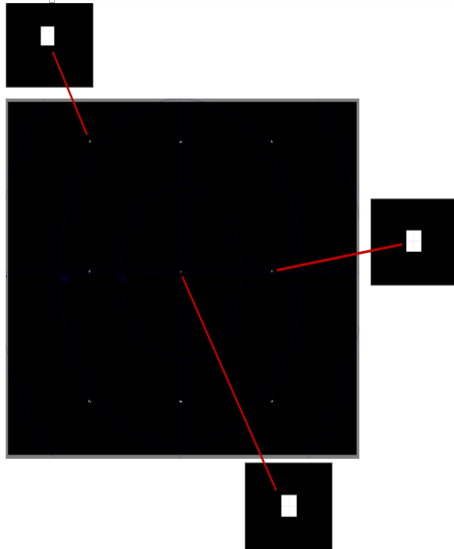


DDFacet(courtesy: C. Tasse, Observatoire de Paris)

Model image, nominal PSF



Model image, pseudo PSF



C'est la fin!!! Merci!!!

Thank you!

