

INTRODUCTION TO QUANTUM FIELD THEORY

(A TOUR DE FORCE)

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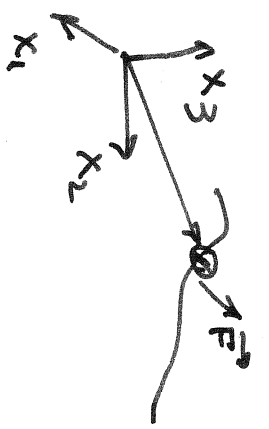
LECTURE I: Fast Review of Fundamental Concepts. ①

Classical Mechanics: $\vec{F} = \frac{d}{dt} \vec{p}$; $\vec{p} = m \frac{d\vec{r}}{dt}$ (E.1)

Constraints, e.g.  $x^2(t) + y^2(t) = l^2$
= constant

(*) Replace Cartesian coordinates with Generalized Coordinates $q_i(t)$ ($i=1, \dots, N$)
all independent of each other.

(*) Instead of forces use Energies



$$\vec{r} = (x_1, x_2, x_3) \quad (\dot{x}_i \equiv \frac{d}{dt} x_i)$$

$$\vec{F}_i = \frac{d}{dt} p_i = \frac{d}{dt} (m \dot{x}_i) = \frac{d}{dt} \frac{\partial}{\partial \dot{x}_i} \left(\frac{1}{2} m \dot{x}_i^2 \right)$$

$$= \frac{d}{dt} \frac{\partial T}{\partial \dot{x}_i}$$

Kinetic Energy T

In Cartesian coordinates $\frac{\partial T}{\partial x_i} = 0$

Conservative systems $\frac{\partial T}{\partial x_i} = 0$

V: potential energy

$$\frac{\partial}{\partial x_i} (T - V) = \frac{d}{dt} \frac{\partial}{\partial \dot{x}_i} (T - V)$$

L (Lagrangian)

Lagrange's Equations

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{x}_i} - \frac{\partial L}{\partial x_i} = 0 \quad (\text{I.2})$$

$$x_i(t) \Rightarrow q_i(t) \quad i=1,2,\dots,N$$

N: number of "degrees of freedom".

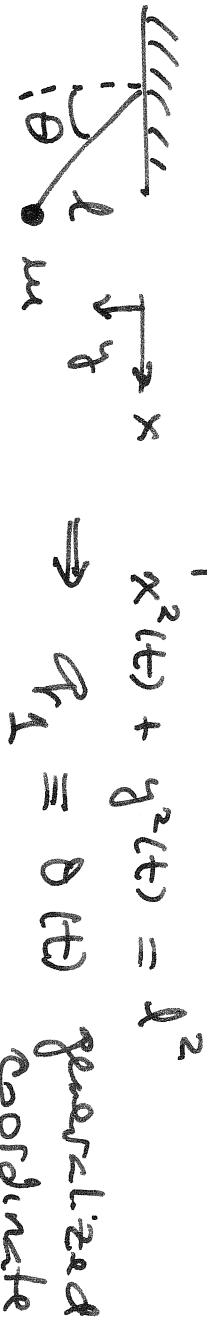
Generalized momentum: $p_i = \frac{\partial L}{\partial \dot{q}_i}$

$$(I.3) \quad \left\{ \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} - \frac{\partial L}{\partial q_i} = 0 \Rightarrow \frac{d}{dt} p_i - \frac{\partial L}{\partial q_i} = 0 \right. \quad (I.4)$$

If $\frac{\partial L}{\partial q^*} = 0 \Rightarrow \frac{d}{dt} p^* = 0 \dots p^* = \text{constant of motion.}$

CONFIGURATION SPACE

The space generated by the generalized coordinates. Example



Configuration space:

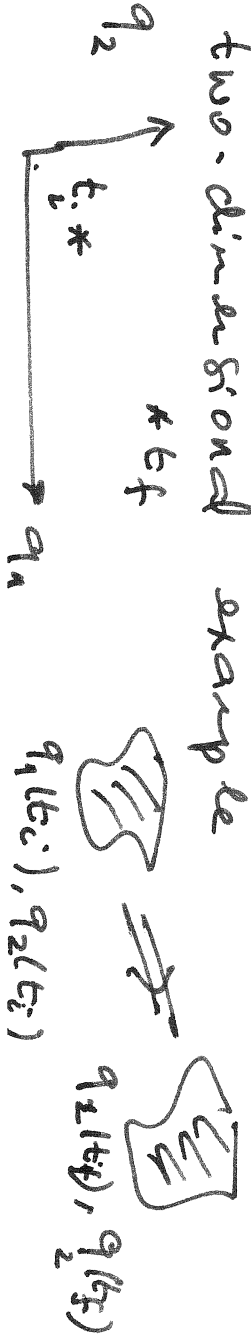


ACTION

$$S = \int_{t_1}^{t_2} L(q_i(t), \dot{q}_i(t), t) dt \quad (I.5)$$

Dim [S] = Energy * Time (e.g. Joule * sec)

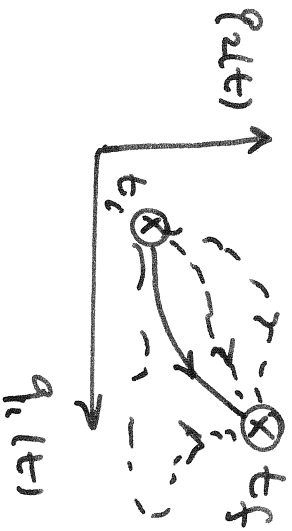
A two-dimensional example



Question: Which

path? only one. Which one? 3

The one that "minimizes" δS



Hamilton Principle

$\delta S = 0$ (I.6) with t_1 & t_2 fixed

$$\delta S = \int_{t_1}^{t_2} \delta L dt = \int_{t_1}^{t_2} \left[\frac{\partial L}{\partial q_i} \delta q_i + \frac{\partial L}{\partial \dot{q}_i} \delta \dot{q}_i \right] dt$$

$$\int_{t_1}^{t_2} dt \frac{\partial L}{\partial q_i} \delta q_i = \int_{t_1}^{t_2} dt \frac{\partial L}{\partial \dot{q}_i} \frac{d}{dt} \delta q_i = \left[\delta q_i \frac{\partial L}{\partial \dot{q}_i} \right]_{t_1}^{t_2} - \underbrace{\delta q_i(t_1)}_{\delta q_i(t_1) = \delta q_i(t_2) = 0}$$

$$- \int_{t_1}^{t_2} dt \delta q_i \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} \right)$$

$$\delta S = \int_{t_1}^{t_2} dt \left[\frac{\partial L}{\partial q_i} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} \right) \right] \delta q_i$$

Using (I.6) & since δq_i is arbitrary,

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} \right) - \frac{\partial L}{\partial q_i} = 0$$

which are Lagrange's eqns (I.3) - (I.4).

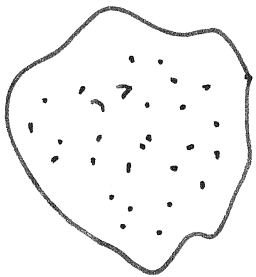
Hamiltonian:

$$H = \sum_{i=1}^N (p_i \dot{q}_i - L) \quad \text{(I.7)}$$

In most cases $H = T + V$, i.e. the total energy of the system.

FIELDS

(4)



$q_i(t)$, $i=1, 2, \dots, N$.

For large N , e.g. $N \approx N_A \approx 6 \times 10^{23}$

treat system as a "continuum"

$q_i(t)$

$i=1, 2, \dots, N \rightarrow \infty$ & non-denumerable,
meaning i becomes a continuous
variable, i.e. $\vec{x} = (x_1, x_2, x_3)$

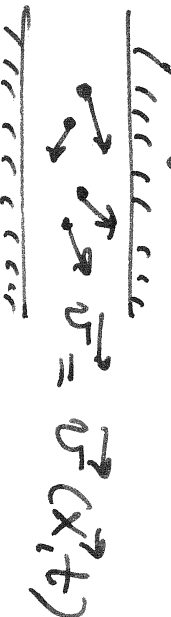
$$q_i(t) \Rightarrow q(\vec{x}, t) \equiv \underbrace{\varphi(\vec{x}, t)}_{\text{FIELD}} \quad (\text{I.8})$$

FIELD: function of \vec{x} & t describing
a continuum \mathcal{V}

Example: i) Deformations of a non-rigid
body



ii) Velocity field in fluid flow:



LAGRANGIAN DENSITY

$$h(q_i(t), \dot{q}_i(t)) \Rightarrow \mathcal{L}(\varphi(\vec{x}, t), \vec{v} \varphi(\vec{x}, t), \frac{\partial \varphi(\vec{x}, t)}{\partial t})$$

$$L \Rightarrow \int d^3x \mathcal{L} \quad (I.9)$$

Relativistic Notation:

$$x^\mu = (x^0 = ct, \vec{x}); \quad x_\mu = (ct, -\vec{x})$$

$$x_\mu = g_{\mu\nu} x^\nu; \quad g_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} = \eta^{\mu\nu}$$

Minkowski space

Metric Tensor

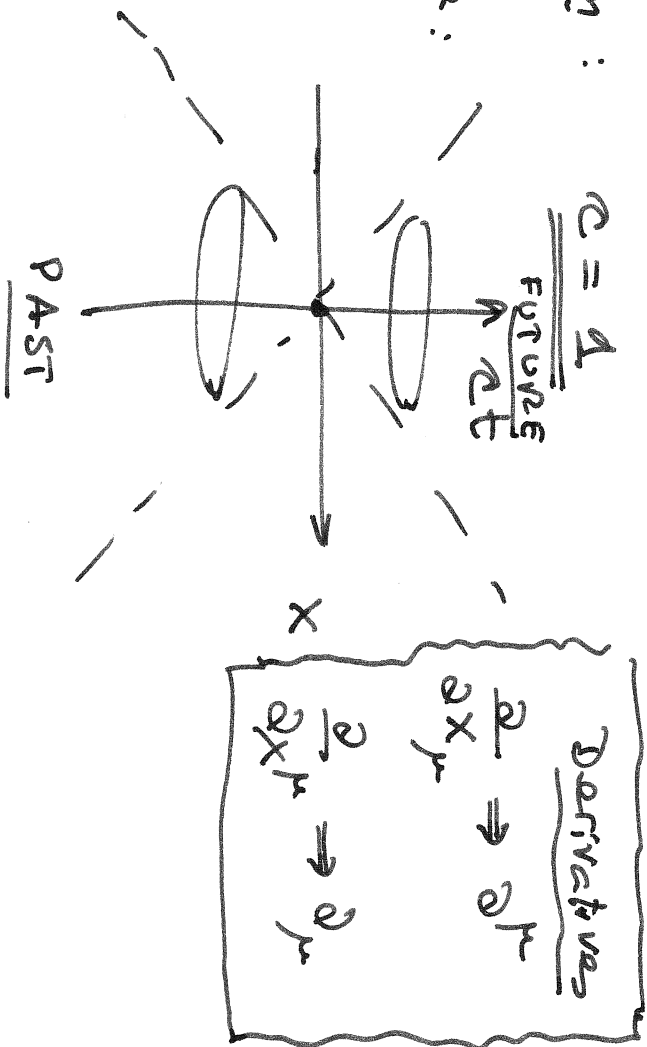
(\vec{x}, t) 4-dimensional space

Distance: $x^\mu x_\mu = x^0^2 - \vec{x}^2$

Convention: $c = 1$

FUTURE
↑
ct

Light Cone:



$$\mathcal{L} = \mathcal{L}(\varphi(x), \partial^\mu \varphi(x)) \quad (I.10)$$

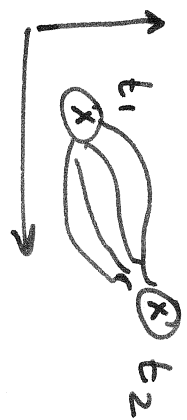
$$S = \int d^4x \mathcal{L}(\varphi(x), \partial^\mu \varphi(x)) \quad (I.11)$$

Hamilton's principle: $\delta S = 0$

Leads to Lagrange's equations for $\varphi(x)$

$\delta S'$: a variation in (1) time & (2) space

i.e. on a surface (3-dimensional)



$$\delta \varphi(t_1) = \delta \varphi(t_2) = 0$$

fixed times.



$$\delta^1 = \int_{t_1}^{t_2} \int_{\Sigma} d^3x \mathcal{L}(\varphi, \partial_r \varphi)$$

time variation: $\varphi \rightarrow \varphi' = \varphi + \delta_0 \varphi$

$$\delta S^1 = \int_{t_1}^{t_2} \int_{\Sigma} d^3x \left[\frac{\partial \mathcal{L}}{\partial \varphi(x)} \delta_0 \varphi + \frac{\partial \mathcal{L}}{\partial (\partial_r \varphi)} \delta_0 (\partial_r \varphi) \right]$$

$$\int_{t_1}^{t_2} dt \int_{\Sigma} d^3x \frac{\partial \mathcal{L}}{\partial (\partial_r \varphi)} \delta_0 (\partial_r \varphi) =$$

$$= \int_{t_1}^{t_2} dt \int_{\Sigma} d^3x \frac{\partial \mathcal{L}}{\partial (\partial_r \varphi)} \partial_r (\delta_0 \varphi)$$

$$= \underbrace{\left[\int_{\Sigma} \delta_0 \varphi \frac{\partial \mathcal{L}}{\partial (\partial_r \varphi)} \right]_{t_1}^{t_2}}_{=0} - \int_{t_1}^{t_2} dt \int_{\Sigma} d^3x \partial_r \frac{\partial \mathcal{L}}{\partial (\partial_r \varphi)} \delta_0 \varphi$$

$$\delta S^1 = \int d^4x \left[\frac{\partial \mathcal{L}}{\partial \varphi(x)} - \partial_r \frac{\partial \mathcal{L}}{\partial (\partial_r \varphi)} \right] \delta_0 \varphi = 0 \quad (\text{I.12})$$

$$\Rightarrow \left[\frac{\partial \mathcal{L}}{\partial \varphi(x)} - \partial_r \frac{\partial \mathcal{L}}{\partial (\partial_r \varphi)} \right] = 0 \quad (\text{I.13})$$

Field Momentum:

$$\left\{ \pi = \frac{\partial \mathcal{L}}{\partial(\partial_0 \varphi)} \right. \quad (\text{I.14})$$

(to mimic $p_i \equiv \partial L / \partial \dot{q}_i$)

Hamiltonian Density:

$$H = \sum_i p_i \dot{q}_i - L \Rightarrow \boxed{H = \sum_i \pi_i \frac{\partial \varphi_i}{\partial t} - \mathcal{L}_{\text{int}}}$$

— x —

Electromagnetic Field (Classical)

The only field in Nature that exists at the classical (i.e. non-quantum) level. Quantum fields associated with leptons (e, μ , τ) & quarks (u, d, s, c, b, t) do not have a classical limit. They simply do not exist at the classical level?

E.M. Field \Rightarrow Maxwell

$$\left\{ \begin{array}{l} \vec{E}(x) \equiv \vec{E}(\vec{x}, t) \\ \vec{B}(x) \equiv \vec{B}(\vec{x}, t) \end{array} \right. \quad (\text{I.16})$$

$$\left. \begin{aligned} \vec{\nabla} \cdot \vec{E}(\vec{x}, t) &= \rho(\vec{x}, t) \\ \vec{\nabla} \times \vec{B}(\vec{x}, t) - \frac{\partial \vec{E}(\vec{x}, t)}{\partial t} &= \vec{J}(\vec{x}, t) \end{aligned} \right\} \text{(I.17)}$$

$$\left. \begin{aligned} \vec{\nabla} \cdot \vec{B}(\vec{x}, t) &= 0 \\ \vec{\nabla} \times \vec{E}(\vec{x}, t) + \frac{\partial \vec{B}(\vec{x}, t)}{\partial t} &= 0 \end{aligned} \right\}$$

Potentials: $\Phi(\vec{x}, t)$ & $\vec{A}(\vec{x}, t) \Rightarrow A^\mu(x)$ (I.18)

$$A^\mu(x) \Rightarrow [\Phi(x), \vec{A}(x)] \quad \text{(I.19)}$$

$$\vec{E} = -\vec{\nabla}\Phi - \frac{\partial \vec{A}}{\partial t}; \quad \vec{B} = \vec{\nabla} \times \vec{A} \quad \text{(I.20)}$$

EM Tensor: $F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu = -F^{\nu\mu}$ (I.21)

$$F^{\mu\nu} = \begin{pmatrix} 0 & -E_1 & -E_2 & -E_3 \\ E_1 & 0 & -B_3 & B_2 \\ E_2 & B_3 & 0 & -B_1 \\ E_3 & -B_2 & B_1 & 0 \end{pmatrix} \quad \text{(I.22)}$$

show: $\mathcal{L}_{EM} = -\frac{1}{4} F^{\mu\nu} F_{\mu\nu}$ (I.23)

in the absence of sources, i.e. in free space ($\rho=0, \vec{J}=0$)... \mathcal{L}_{EM} takes form

$$\left[\partial^\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} - \frac{\partial \mathcal{L}}{\partial (A_\nu)} = 0 \right] \quad \text{(I.24)}$$

Show:

$$\partial_\mu F^{\mu\nu}(x) = J^\nu(x) \quad (\text{I.25}) \quad (9)$$

$$\partial_\lambda F_{\mu\sigma} + \partial_\mu F_{\nu\lambda} + \partial_\nu F_{\lambda\mu} = 0 \quad (\text{I.26})$$

Show that Eq. (I.25) leads to the two inhomogeneous Maxwell eqns. and (I.26) leads to the other two (homogeneous) eqns. -

HINT: $\frac{\partial \mathcal{L}}{\partial(\partial^\beta A^\alpha)}$ leads to 4 identical terms

$$-\frac{1}{4} g_{\lambda\sigma} g_{\nu\tau} \left\{ \delta_\beta^\mu \delta_\alpha^\sigma F^{\lambda\nu}; \delta_\beta^\nu \delta_\alpha^\sigma F^{\lambda\mu}; \delta_\beta^\lambda \delta_\alpha^\sigma F^{\mu\nu}; \delta_\beta^\sigma \delta_\alpha^\lambda F^{\mu\nu} \right\} \quad (\text{I.27})$$

hence

$$\frac{\partial \mathcal{L}}{\partial(\partial^\beta A^\alpha)} = -g_{\lambda\beta} g_{\nu\alpha} F^{\lambda\nu} = -\bar{F}_{\beta\alpha} = \bar{F}_{\alpha\beta} \quad (\text{I.28})$$

We shall return later to the EM field to discuss its quantization.

Noether's Theorem

For every continuous symmetry of a Lagrangian there is a conserved current ($\partial_\mu T_\nu(x) = 0$) built from the fields in the Lagrangian.

Proof Assume $\mathcal{L} = \mathcal{L}(\varphi(x), \partial_\mu \varphi(x))$

trivially generalized to an \mathcal{L} depending on more than one field.

Perform a transformation of the field

$$\varphi(x) \longrightarrow \varphi'(x) = \varphi(x) + \delta\varphi(x) \quad (\text{I.29})$$

e.g. $\varphi(x) \rightarrow \varphi'(x) = e^{i\theta} \varphi(x) \simeq (1 + i\theta + \dots)\varphi(x)$

The change in \mathcal{L} is

$$\delta\mathcal{L} = \frac{\partial\mathcal{L}}{\partial\varphi} \delta\varphi + \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi)} \delta(\partial_\mu\varphi) \quad (\text{I.30})$$

Use Lagrange's eqns $\frac{\partial\mathcal{L}}{\partial\varphi} = \partial_\mu \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi)}$

to write

$$\delta\mathcal{L} = \partial_\mu \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi)} \delta\varphi + \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi)} \delta\varphi$$

$$\boxed{\delta\mathcal{L} = \partial_\mu \left[\frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi)} \delta\varphi \right]} \quad (\text{I.31})$$

If the transformation (I.29) leaves the Lagrangian invariant, i.e.

if $\delta \mathcal{L} = 0$ then $T_\mu \equiv \frac{\partial \mathcal{L}}{\partial(\partial_\mu \varphi)}$ (I.32)

is conserved, i.e.

$\partial^\mu T_\mu = 0$ (I.33)

Example:

Classical Electrodynamics in vacuum

$\mathcal{L}_{CED} = -\frac{1}{4} F_{\mu\nu}(x) F^{\mu\nu}(x)$
 $= -\frac{1}{4} (\partial_\mu A_\nu(x) - \partial_\nu A_\mu(x))$
 $\times (\partial^\mu A^\nu(x) - \partial^\nu A^\mu(x))$

$A_\mu(x) \Rightarrow A'_\mu(x) = A_\mu(x) + \partial_\mu \Lambda(x)$

$\Rightarrow \mathcal{L}'_{CED} = \mathcal{L}_{CED}$

$T^\mu = \frac{\partial \mathcal{L}}{\partial(\partial_\mu A_\nu)} = -\frac{1}{4} (\partial^\mu A^\nu - \partial^\nu A^\mu)$

$\partial_\mu T^\mu = -\frac{1}{4} (\partial_\mu \partial^\mu A^\nu - \partial^\nu \partial_\mu A^\mu) = 0.$

from equation (Lorentz) of motion: $\square^2 A^\nu = 0$ in free space

(Non-relativistic) QUANTUM MECHANICS

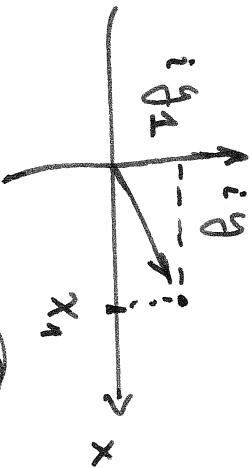
(12)

Schrodinger equation: (Atomic, Molecular, Physics & CHEMISTRY)

$$(I.34) \quad \sqrt{-1} \hbar \frac{\partial \psi(\vec{x}, t)}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi(\vec{x}, t) + V \psi(\vec{x}, t)$$

Complex numbers: $Z = (z_1, z_2)$ a pair of

numbers, e.g. $z_1 = x$, $z_2 = \sqrt{-1} y \equiv iy$



$$\boxed{i^2 = -1} \quad (I.35)$$

$$\boxed{e^{i\pi} = -1} \quad (I.36)$$

i RELATES e WITH π ∇
Profound meaning.

— x —

$\psi(\vec{x}, t)$ a complex function

$$|\psi(\vec{x}, t)|^2 = \psi(\vec{x}, t) \psi^*(\vec{x}, t)$$

probability of finding the "particle"
described by $\psi(\vec{x}, t)$ at point \vec{x}
& at time t .

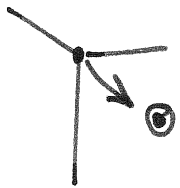
"WAVE-PARTICLE DUALITY"

A non statistical statement

CLASSICAL UNIVERSE

(13)

"particle":



localized point object
lang object whose dimensions
are negligible in the
problem at hand)

"wave"



delocalized object

Opposite concepts ONLY in the electrical domain. At the quantum level, this concept/definition breaks down.

We should call an e^- neither a particle nor a wave, but something else, e.g. "dualon".

Explanation of this conundrum is not to be found in QUANTUM MECHANICS but only in QUANTUM FIELD THEORY.

e.g. e^- . What is the electron?

There is a quantum field describing the electron (Dirac field). Its quantum excitations (of the field) manifest themselves as particles.

Schrödinger current:

Assume $V=0$ & write the eqn. and its complex conjugate

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi \quad \& \text{ multiply by } \psi^*$$

$$-i\hbar \frac{\partial \psi^*}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi^* \quad \& \text{ multiply by } \psi$$

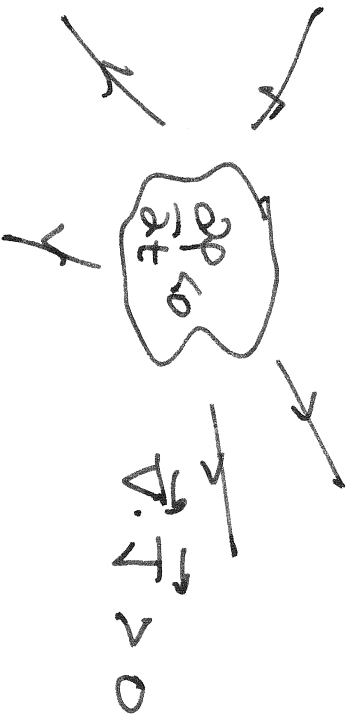
Subtracting gives

$$\frac{\partial}{\partial t} (\psi^* \psi) = \frac{\partial}{\partial t} (\psi^* \nabla^2 \psi - \psi \nabla^2 \psi^*)$$

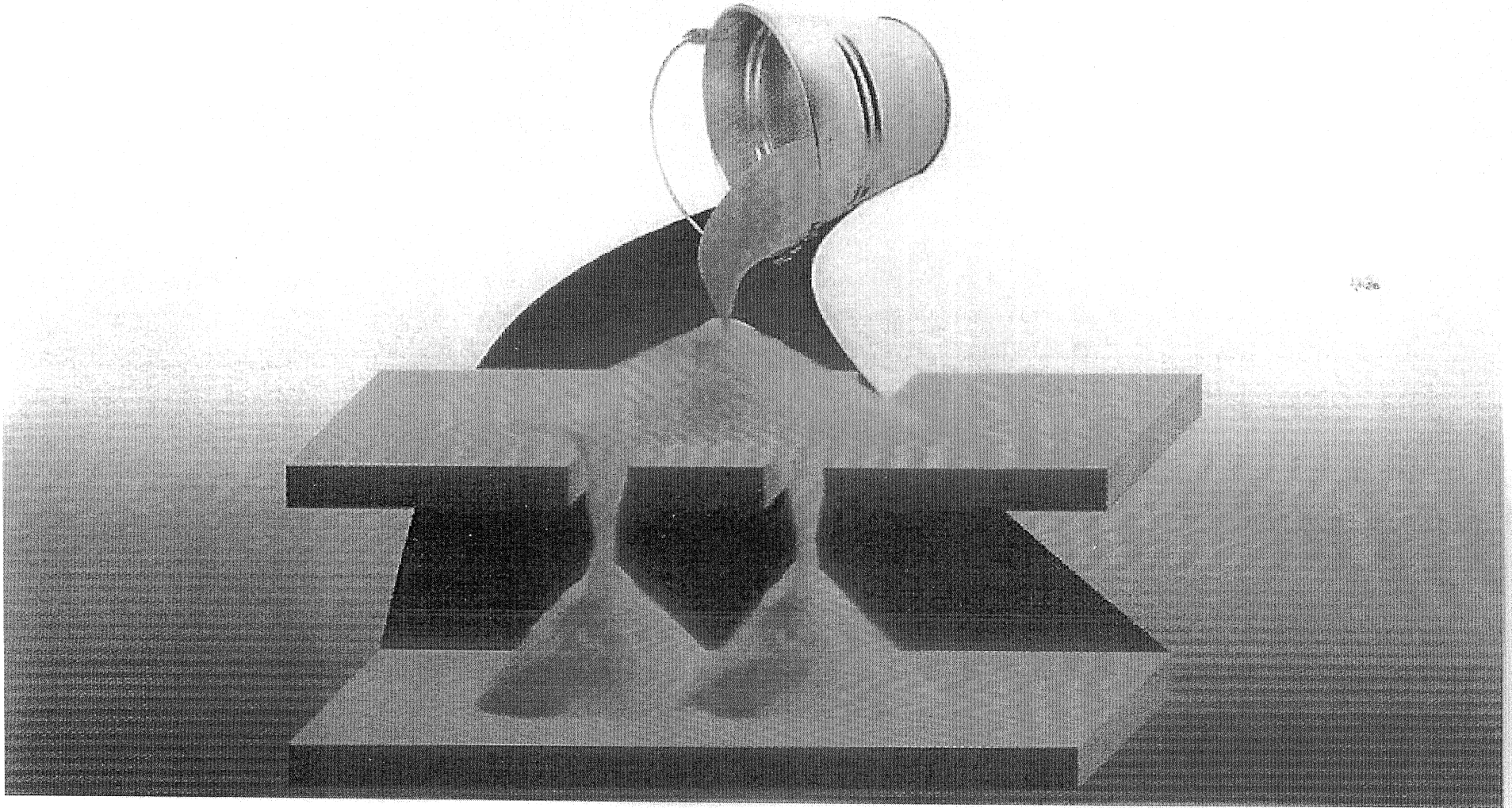
$$\frac{\partial}{\partial t} |\psi(\vec{x}, t)|^2 = \frac{\partial \rho(\vec{x}, t)}{\partial t} = -\vec{\nabla} \cdot \vec{J}(\vec{x}, t) \quad (E.37)$$

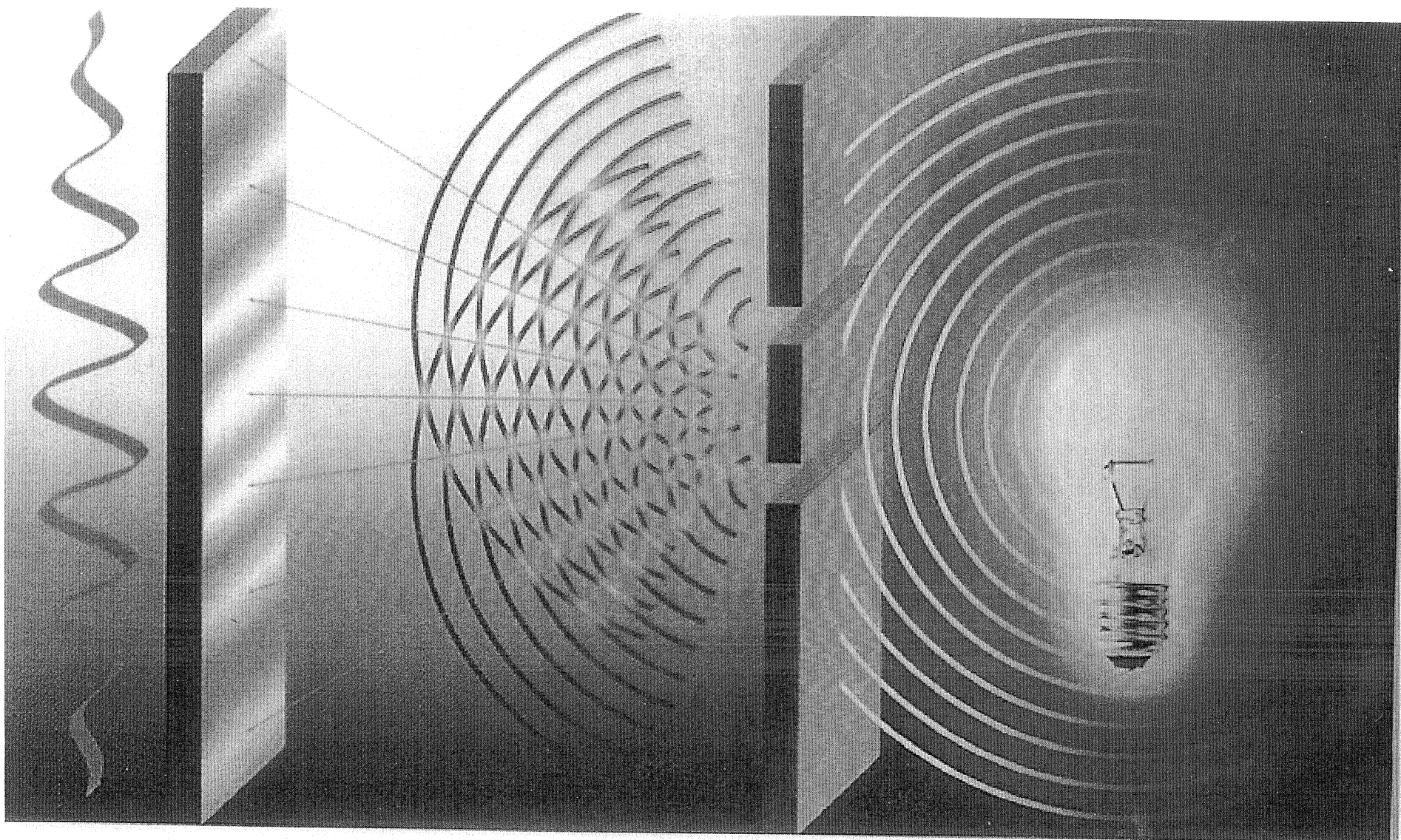
$$\vec{J}(\vec{x}, t) = \frac{i\hbar}{2m} (\psi \vec{\nabla} \psi^* - \psi^* \vec{\nabla} \psi) \quad (E.38)$$

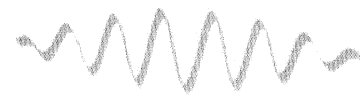
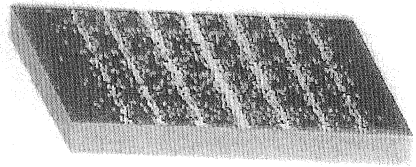
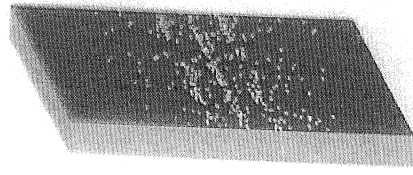
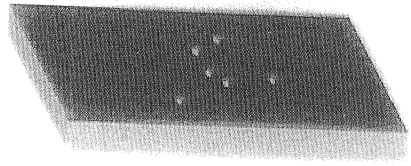
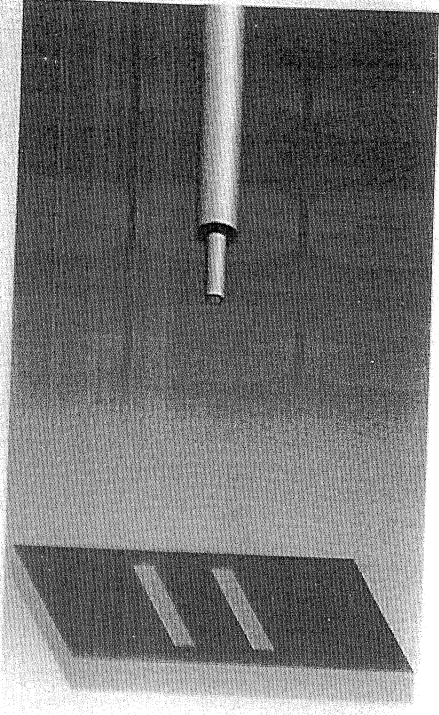
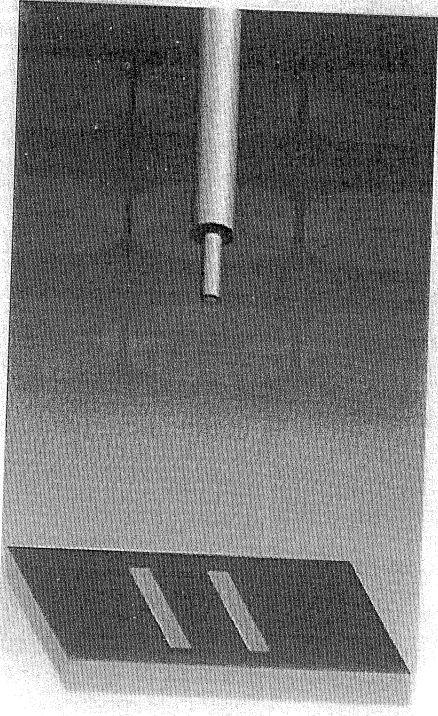
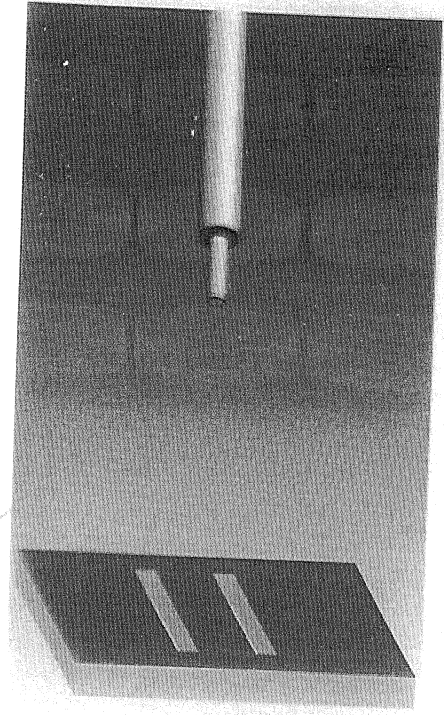
If $\frac{\partial \rho}{\partial t} < 0$ in a volume then $\vec{\nabla} \cdot \vec{J} > 0$

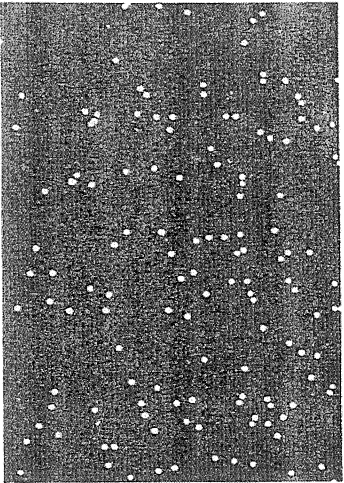


SLIDES: Classical & Quantum interference.

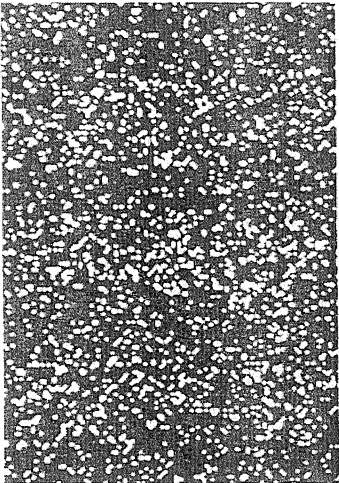




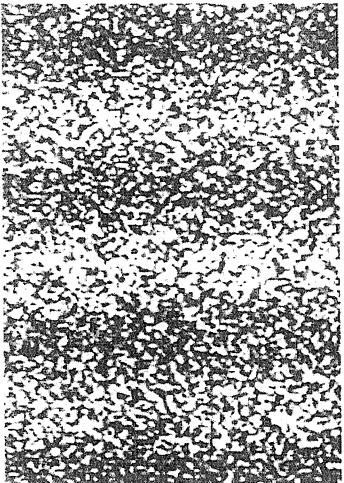




(b) After 100 electrons



(c) After 3000 electrons

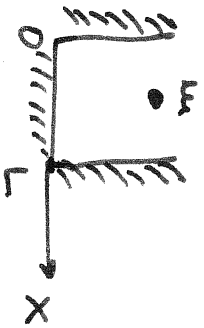


70,000

TUTORIAL I

(19)

1) Solve the Schrödinger equation for a particle in a one-dimensional box (line) of length L , bounded by infinite walls. Consider both $E < 0$ and $E > 0$. Consider the time-independent case.



$$V(x) = \begin{cases} \infty & x < 0 \\ 0 & 0 < x < L \\ \infty & x > L \end{cases}$$

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi(x)}{dx^2} + V(x) \psi(x) = E \psi(x)$$

a) $E < 0$: there are no solutions.

b) $E > 0$

$$\psi(x) \equiv u(x) \quad k^2 \equiv \frac{2mE}{\hbar^2} > 0$$

$$\frac{d^2 u(x)}{dx^2} + k^2 u(x) = 0$$

$$u(x)|_{x=0} = 0 \quad u(x)|_{x=L} = 0$$

$$u(x) = A \sin kx + B \cos kx$$

$$\text{Let } B = 0 \quad \& \quad kL = n\pi$$

$$\therefore \left\{ u(x) = A \sin\left(\frac{n\pi}{L}x\right) \right\}$$

$$E_n = \frac{\hbar^2}{2m} k_n^2 = \frac{\hbar^2}{2m} \frac{n^2 \pi^2}{L^2}$$

$$\boxed{E_n = \frac{\hbar^2}{2m} \frac{n^2 \pi^2}{L^2}}$$

$$E_n \propto n^2$$